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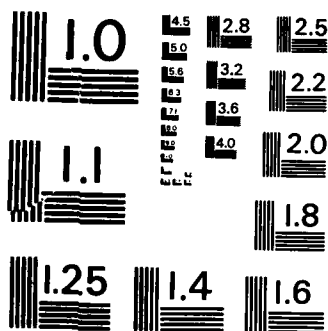   MATHEMATICS RESEARCH CENTER P CONSTANTIN NOV 85
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NOTE ON LOSS OF REGULARITY FOR SOLUTIONS
OF THE 3-D INCOMPRESSIBLE EULER
AND RELATED EQUATIONS

Petre Constantin

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NOTE ON LOSS OF REGULARITY FOR SOLUTIONS OF THE 3-D
INCOMPRESSIBLE EULER AND RELATED EQUATIONS

Petre Constantin[*]
Technical Summary Report #2884
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ABSTRACT

We prove that as long as the solution to the ideal flow equation in three

space dimensions is smooth the solutions of slightly viscous incompressible

Navier-Stokes equations are smooth. We describe blow-up results for a semi-

Lagrangian short time approximation of the incompressible Euler equation.

AMS (MOS) Subject Classifications: 35Q10, 35Q99, 76C99, 76D05

Key Words: blow-up, Riesz transforms

Work Unit Number 1 (Applied Analysis)

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SIGNIFICANCE AND EXPLANATION

One of the central problems in the mathematical theory of turbulence is that of breakdown of smooth (indefinitely differentiable) solutions to the equations of motion. In 1934 J. Leray advanced the idea that turbulence may be related to the spontaneous appearance of singularities in solutions of the 3-D incompressible Navier-Stokes equations. The problem is still open. We ~~show in~~ this report ^{shows} that breakdown of smooth solutions to the 3-D incompressible slightly viscous (i.e. corresponding to high Reynolds numbers, or "highly turbulent") Navier-Stokes equations cannot occur without breakdown in the corresponding solution of the incompressible Euler (ideal fluid) equation. ~~We~~ prove ^{It is} then that solutions of distorted Euler equations, which are equations closely related to the Euler equations for short term intervals, do breakdown. *Keywords:* ...

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NOTE ON LOSS OF REGULARITY FOR SOLUTIONS OF THE 3-D

INCOMPRESSIBLE EULER AND RELATED EQUATIONS

Petre Constantin[*]

Introduction

The purpose of this paper is twofold: first to discuss the relationship between the breakdown of smooth solutions to incompressible three-dimensional Euler and Navier-Stokes equations; and secondly to present blow-up results for distorted Euler equations.

Both the Navier-Stokes equations and the Euler equations possess local (in time) smooth solutions. Moreover, as the viscosity vanishes the solutions to the Navier-Stokes equations converge uniformly on a short time interval to the solution of the Euler equation ([5],[7]). Adapting the method of Kato ([5]) and using a very simple ODE lemma we prove in Section 1 that as long as the solution to the Euler equation is smooth the solutions to slightly viscous Navier-Stokes equations with the same initial data are smooth.

Sections 2 and 3 are devoted to blow-up results for distorted Euler equations. Differentiating the Euler equations one obtains a quadratic equation for the Jacobian matrix of the velocity vector:

$$\frac{\partial U}{\partial t} + (u \cdot \nabla)U + U^2 = P$$

where u is the velocity vector, $U = (\frac{\partial u_i}{\partial x_j})i,j$ and $P = (\frac{\partial^2 p}{\partial x_i \partial x_j})$ with p the pressure.

One can use the incompressibility condition $TrU = 0$ to express P in terms of U. Passing to Lagrangian coordinates, the differentiated Euler equations become

$$(0.1) \qquad \frac{\partial U}{\partial t} + U^2 + R(t)(TrU^2) = 0$$

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where $R(t)$ is a matrix of singular integral operators with time varying kernels. What we call the distorted Euler equations are obtained from the above form of the genuine Euler equations by replacing $R(t)$ by $R(0)$:

$$(0.2) \qquad \frac{\partial U}{\partial t} + U^2 + R(0)(TrU^2) = 0 \ .$$

Although these equations are good short time approximations of the Euler equations, the blow-up arguments have no direct bearing on the Euler equations.

In Section 2 we discuss the periodic case and we show, by a localizati argument reminiscent of the one in [2], that a large class of initial data lead to kdown of the solution of (0.2). The conditions on the initial data do not involve any ness assumption but exclude Jacobians. Another drawback in the periodic case i e fact that incompressibility, $TrU = 0$, is not preserved. This fact is due to the non nishing of the mean of TrU^2 but it is not the major reason for the blow-up. (One can m dify slightly the equations in order to preserve the constraint $TrU = 0$ and still prove breakdown.) Moreover, in the whole space case the equations (0.2) do preserve incompressibility. Section 3 treats solutions of the distorted Euler equations in the whole space. C. Foias found ([4]) that if the initial data for (0.2) have the form

$$(0.3) \qquad U_0(x) = \beta_0(|x|)(I - n\pi(x))$$

where n is the dimension and $\pi(x) = \left(\frac{x_i x_j}{|x|^2}\right) i,j = 1,\ldots,n,$ then this form is retained by the solution $U(t,x)$ of (0.2) and leads to a simple equation for the scalar quantity β. We generalize slightly his result by allowing U_0 to possess an antisymmetric part, corresponding to the vorticity. We obtain a system of integro-differential equations for two scalar quantities $\beta(t,r)$ (corresponding to the size of the deformation tensor) and $\gamma(t,r)$ corresponding to the modulus of the vorticity. For $\gamma \equiv 0$ we recover the Foias equation. The success of the reduction in the number of variables and unknowns is due to a covariance property of equation (0.2) with respect to an action of $O(n)$. We prove breakdown for solutions starting from initial data of the special form

$$U_0(x) = \beta_0(|x|)(I - 3\pi(x)) + \gamma_0(|x|) \frac{x}{|x|} \times \ .$$

If one takes the antisymmetric part of the three-dimensional distorted Euler equations and if one identifies 3×3 antisymmetric matrices J with the vectors given by $J = \omega \times$ one obtains the equation

(0.4)
$$\frac{\partial \omega}{\partial t} = U\omega$$

which is the analogue of the vorticity equation for incompressible Euler flows.

In [3] the simple one-dimensional model equation for the three-dimensional vorticity equation

(0.5)
$$\frac{\partial \omega}{\partial t} = \omega H \omega \qquad (H = \text{Hilbert transform})$$

was suggested. The breakdown of solutions to (0.2) is very similar to that of solutions to (0.5): The quantity corresponding to the deformation tensor (the symmetric part of U in the case of (0.2), $H\omega$ in the case of (0.5)) becomes infinite in regions when the quantity corresponding to the vorticity (denoted ω in both cases) is zero.

It is a pleasure to thank A. Majda for suggesting the result of Section 1 and for stimulating discussions. I had many interesting conversations with C. Foias and S. Klainerman. It is an equal pleasure to thank them both.

1. A comparison result

Let us consider a solution v of the incompressible Euler equations

(1.1)
$$\begin{cases} \dfrac{\partial v}{\partial t} + (v \cdot \nabla)v = \nabla p + f \\[2mm] \text{div } v = 0 \\[2mm] v(0, \cdot) = v_0 \end{cases}$$

in either $\mathbb{R}^3$ or T^3 (the three-dimensional torus). In this section we prove that as long as $v(t, \cdot)$ is smooth, the solutions to slightly viscous incompressible Navier-Stokes equations having v_0 as initial data are smooth.

We use the notation H^m for the Sobolev spaces $H^m = H^m(\mathbb{R}^3)$ (resp. $H^m = H^m(T^3)$) and $(\cdot, \cdot)_m$, $|\cdot|_m$ for the corresponding scalar products and norms.

Theorem 1.1

Let $v = v(t, x)$ be a solution of (1.1) for $0 < t < T$, satisfying

(1.2)
$$|v_0|_{m+2} < \infty \quad \text{for some} \quad m > 3 \; ,$$

(1.3)
$$\int_0^T |\nabla \times v|_{L^\infty} \, dt < \infty \; .$$

Then there exists $\nu_0 = \nu_0(T; |v_0|_{m+2}, \int_0^T |\nabla \times v|_{L^\infty} \, dt)$ such that, for every $0 < \nu < \nu_0$ the solution to the Navier-Stokes equation

(1.4)
$$\begin{cases} \dfrac{\partial u}{\partial t} + (u \cdot \nabla)u = \nu \Delta u + \nabla q + f \\[2mm] \text{div } u = 0 \\[2mm] u(0, \cdot) = v_0 \end{cases}$$

is smooth on $[0, T]$. More precisely

(1.5)
$$\sup_{t \in [0,T]} |u(t) - v(t)|_m < \nu \gamma_m$$

for some γ_m depending on T, $|v_0|_{m+2}$, $\int_0^T |\nabla \times v|_{L^\infty} \, dt$.

Let us emphasize here that T is not assumed to be small. Instead it is assumed that $v(t)$ belongs to H^{m+2} for $t \in [0, T]$. Indeed, assumption (1.3) was proven by

-4-

J. T. Beale, T. Kato and A. Majda ([1]) to be a sufficient condition for higher regularity. Their result can be stated as follows

Theorem 1.2 (Beale, Kato, Majda)

Assume $\int_0^T |\nabla \times v|_{L^\infty} \, dt < \infty$. Let $s > 3$, $v_0 \in H^s$. There exists a constant c depending on T, s, $\int_0^T |\nabla \times v|_{L^\infty} \, dt$, $\|v_0\|_s$, such that

(1.6)
$$\|v(t)\|_s < c \quad \text{for} \quad t < T .$$

In order to prove Theorem 1.1 let us consider the difference $w = u - v$. Then w will satisfy

(1.7)
$$\begin{cases} \dfrac{\partial w}{\partial t} - \nu \Delta w + (v \cdot \nabla)w + (w \cdot \nabla)v + (w \cdot \nabla)w = \nu \Delta v + \nabla r \\[2mm] \text{div } w = 0 \\[2mm] w(0, \cdot) = 0 . \end{cases}$$

We take the scalar product of (1.7) with w in H^m, $m > 3$ and use

(1.8)
$$\left| ((v \cdot \nabla)w, w)_m \right| < c_m \|v\|_m \|w\|_m^2 ,$$

(1.9)
$$\left| (w \cdot \nabla v, w)_m \right| < c_m \|v\|_{m+1} \|w\|_m^2$$

for $v \in H^{m+1}$, div $v = 0$, $w \in H^{m+1}$, div $w = 0$ (see [5]).

Using the fact that $-\nu(\Delta w, w)_m > 0$ we obtain

(1.10)
$$\frac{d}{dt} \|w\|_m < \nu \|\Delta v\|_m + c_m \|v\|_{m+1} \|w\|_m + c_m \|w\|_m^2 .$$

Let us multiply (1.10) by $\exp(-c_m \int_0^t \|v\|_{m+1} ds)$ and consider the quantity $y = \|w\|_m \exp(-c_m \int_0^t \|v\|_{m+1} ds)$. We obtain the inequality

(1.11)
$$\begin{cases} \dfrac{dy}{dt} < \nu F(t) + G y^2 \\[2mm] y(0) = 0 \end{cases}$$

with

(1.12)
$$F(t) = \|\Delta v\|_m \exp - c_m \int_0^t \|v\|_{m+1} ds$$

and

(1.13)
$$G = c_m \exp c_m \int_0^T \|v\|_{m+1} dt .$$

-5-

We shall make use now of an elementary lemma:

Lemma 1.3

Let $T > 0$, $G > 0$ be given constants and let $F(t)$ be a nonnegative continuous function on $[0,T]$. Let ν_0 be defined by

$$(1.14) \qquad \nu_0 = \frac{1}{8TG \int_0^T F(t)\,dt} \, .$$

Then, for every $0 < \nu < \nu_0$, every solution $y \geq 0$ of (1.11) is uniformly bounded on $[0,T]$ and

$$(1.15) \qquad y(t) < \text{Min}\left\{\frac{3}{2TG} \, , \, 12\nu \int_0^T F(t)\,dt\right\} \, .$$

Proof of lemma

Let us define ε by

$$(1.16) \qquad \varepsilon = \text{Min}\left\{\frac{1}{4T^2 G} \, , \, 16\nu^2 G\left(\int_0^T F(t)\,dt\right)^2\right\} \, .$$

We divide (1.11) by $\left(1 + \sqrt{\frac{G}{\varepsilon}}\, y\right)^2$: $\quad \dfrac{\frac{dy}{dt}}{\left(1 + \sqrt{\frac{G}{\varepsilon}}\, y\right)^2} < \nu F + \varepsilon$. We integrate between 0 and t:

$$(1.17) \qquad \sqrt{\frac{\varepsilon}{G}} \, \frac{1}{1 + \sqrt{\frac{G}{\varepsilon}}\, y(t)} > \sqrt{\frac{\varepsilon}{G}} - \varepsilon T - \nu \int_0^T F(t)\,dt \, .$$

The choice $\varepsilon < \dfrac{1}{4T^2 G}$ implies $\varepsilon T < \dfrac{1}{2}\sqrt{\dfrac{\varepsilon}{G}}$ and, for $\nu < \nu_0$ one has

$\nu \int_0^T F(t)\,dt < \dfrac{1}{4}\sqrt{\dfrac{\varepsilon}{G}}$. Indeed, if $\varepsilon = 16\nu^2 G\left(\int_0^T F(t)\,dt\right)^2$ the last inequality is an

equality and if $\varepsilon = \dfrac{1}{4T^2 G}$ it follows from

$$\frac{1}{4}\sqrt{\frac{\varepsilon}{G}} = \frac{1}{8TG} = \nu_0 \int_0^T F(t)\,dt > \nu \int_0^T F(t)\,dt \, .$$

Thus (1.17) becomes $\quad \dfrac{1}{1 + \sqrt{\frac{G}{\varepsilon}}\, y(t)} > \dfrac{1}{4} \quad$ which implies (1.15).

-6-

We return to the proof of Theorem 1.1.

We apply Lemma 1.3 to (1.11) with F,G defined in (1.12), (1.13). We find ν_0 depending on T, m, $\int_0^T |v|_{m+2} dt$ such that, if $0 < \nu < \nu_0$ and as long as $w(t)$ belongs to H^{m+2}, $t < T$, one has

$$(1.18) \qquad\qquad\qquad\qquad |w(t)|_m < \gamma_m \nu$$

for some γ_m depending on T, m, $\int_0^T |v|_{m+2} dt$.

Using standard calculus inequalities one can find bounds of the type

$$(1.19) \qquad |w(t)|_{m+2} < |v(t)|_{m+2} + \left[|v_0|_{m+2} + \int_0^T |f|_{m+2} dt\right] \exp\left(c_m \int_0^T |w|_m + |v|_m \, ds\right) .$$

Since the validity of (1.18) depends upon $w(t)$ belonging to H^{m+2} but not upon the size of $|w(t)|_{m+2}$ one can argue by contradiction and infer that $|w(t)|_{m+2}$ cannot become infinite for $t < T$ and that (1.18) is true for all $t < T$. We omit further details.

-7-

2. Distorted Euler equations

In this section we prove breakdown of smooth solutions of a "semi-Lagrangian" version of the Euler equations. We start by recalling the Euler equation in $\mathbb{R}^n$ or T^n

$$(2.1) \quad \begin{cases} \partial_t u_i + u_j \partial_j u_i = \partial_i p, & j,i = 1,\ldots,n \ ; \\ \partial_i u_i = 0 \\ u(0,\cdot) = u_0(\cdot) \ . \end{cases}$$

(Here $\partial_t = \dfrac{\partial}{\partial t}$, $\partial_j = \dfrac{\partial}{\partial x_j}$ and summation convention is used.)

Differentiating (2.1) we obtain

$$(2.2) \quad \begin{cases} \partial_t U + (u\cdot\nabla)U + U^2 = P \\ \mathrm{Tr}\,U = 0 \\ U(0,\cdot) = U \ , \end{cases}$$

where U is the $n \times n$ matrix $U = (\partial_j u_i)$, $i = 1,\ldots,n$, $j = 1,\ldots,n$ and P is the Hessian of the pressure $P = (\partial^2_{ij} p)$, $i,j = 1,\ldots,n$. The constraint $\mathrm{Tr}\,U = 0$ (incompressibility) is maintained if $\mathrm{Tr}\,P = \mathrm{Tr}\,U^2$. This means that p solves $\Delta p = \mathrm{Tr}\,U^2$ and therefore the matrix P can be expressed in terms of U:

$$(2.3) \quad P = (-R_i R_j (\mathrm{Tr}\,U^2)), \quad i,j = 1,\ldots,n$$

where R_i are the Riesz transforms defined by

$$(2.4) \quad R_i = (-\Delta)^{-1/2} \partial_i \ .$$

Let us denote by R the operator acting on $n \times n$ matrix valued functions $M = (m_{ij})$, $i = 1,\ldots,n$, $j = 1,\ldots,n$ by

$$(2.5) \quad (RM)_{ij} = R_i R_k (m_{kj}) \ .$$

We identify scalar functions f with the matrices $f\cdot I$ where I is the $n \times n$ identity matrix. The (differentiated) Euler equations can be written as

$$(2.6) \quad \begin{cases} \partial_t U + (u\cdot\nabla)U + U^2 + R(\mathrm{Tr}\,U^2) = 0 \\ \mathrm{Tr}\,U = 0 \\ U(0,\cdot) = U_0(\cdot) \ . \end{cases}$$

If one passes to Lagrangian coordinates in (2.6) that is, if one uses the change of variables $\alpha \overset{\phi^t}{\longmapsto} x(t,\alpha)$ for $x(t,\alpha)$ solving

-8-

$$(2.7) \qquad \begin{cases} \dfrac{dx}{dt} = u(t,x) \\[2mm] x(0,\alpha) = \alpha \end{cases}$$

the Euler equations (2.6) become

$$(2.8) \qquad \begin{cases} \partial_t V + V^2 + R(t)(TrV^2) = 0 \\[2mm] TrV = 0 \\[2mm] V(0,\cdot) = V_0 \end{cases}$$

where $V(t,\alpha) = U(t,x(t,\alpha))$ is the pullback of U and $R(t)$ is the pullback of R through ϕ^t:

$$(2.9) \qquad R(t)M = [R(M\cdot(\phi^t)^{-1})]\cdot\phi^t .$$

More precisely if $k_{ij}(x,y)$ is a kernel for R_iR_j, a kernel for $R(t)_{ij}$ will be $k^t_{ij}(\alpha,\beta) = k_{ij}(x(t,\alpha),x(t,\beta))$. (We used the well-known fact that determinant of Jacobian of ϕ^t is one.) At $t = 0$ the operator $R(t)$ coincides with the Riesz operators $R(0) = R$; this because $x(0,\alpha) = \alpha$. The distorted Euler equations are obtained from the genuine Euler equations (2.8) by freezing $R(t)$ at $t = 0$:

$$(2.10) \qquad \begin{cases} U_t + U^2 + R(TrU^2) = 0 \\[2mm] U(0,\cdot) = U_0 . \end{cases}$$

Let us note that while (2.10) are valid approximations of (2.8) for a short time the blow up arguments that we are going to give have no direct bearing on the Euler equations.

The equations (2.10) are well-posed in a variety of spaces. For instance we can consider the Sobolev spaces $(H^s)^{n^2}$ of matrices with entries in H^s, $s > \frac{n}{2}$. If $s > \frac{n}{2}$ H^s are Banach algebras under pointwise multiplication; the operators R_j are bounded in H^s (for any s, of course). We conclude that, if U is a solution of (2.10)

$$\frac{d}{dt} \|U\|_s \leq c_s \|U\|_s^2$$

and the local existence and uniqueness of solutions of (2.10) follow in standard manner.

We shall treat first the periodic case; we seek solutions to (2.10) which satisfy $U(x + Le_i) = U(x)$ for any $e_i = (0,\dots,1,\dots,0)^t$ and some $L > 0$. We may assume $L = 1$ without loss of generality. Alternately, we shall refer to U as being defined on the n dimensional torus $T^n = \mathbf{R}^n/\mathbf{Z}^n$. Let us denote, for a point x in T^n by $\Gamma_i(x)$ the

-9-

i-th principal circle passing through x:

(2.11) $\qquad \Gamma_i(x) = \{y \in T^n | y_j = x_j \bmod 1, \ j = 1,\ldots n, \ j \neq i\}; \ i = 1,\ldots,n$.

Let us denote, for a matrix U, by S and J the symmetric and respectively antisymmetric parts of U:

$$S = \frac{U + U^*}{2}; \quad J = \frac{U - U^*}{2} \quad \text{where } U^* \text{ is the transposed of } U.$$

Theorem 2.1

Let U_0 be a smooth $n \times n$ matrix valued function on T^n satisfying

(i) $\mathrm{Tr}\, U_0(x) = 0$ for all $x \in T^n$.

(ii) There exists $x_0 \in T^n$ and i, $1 < i < n$ such that

$$\mathrm{supp}\, J_0 \cap \Gamma_i(x_0) = \emptyset \quad (J_0 = \frac{U_0 - U_0^*}{2}) \quad \text{and}$$

$$\int_{\Gamma_i(x_0)} U_{0ii} dx_i < 0 \quad \text{(no summation).}$$

Then the solution of (2.10) having U_0 as initial data breaks down in finite time. More precisely the symmetric part of the solution $U(t,x)$ becomes infinite near $\Gamma_i(x_0)$ in finite time.

Proof

Let us introduce first some notation. We denote for two matrices M, N by $(M;N)$ the scalar product

(2.12) $\qquad (M;N) = \mathrm{Tr}\, MN^*$.

For two matrix valued functions on T^n we denote by $\langle M,N \rangle$ the scalar product

(2.13) $\qquad \langle M,N \rangle = \int_{T^n} (M(x);N(x)) dx$.

Let us first remark that the operator R is symmetric:

(2.14) $\qquad \langle RM,N \rangle = \langle M,RN \rangle$.

Let us break (2.10) into its symmetric and antisymmetric parts:

(2.15) $\qquad \partial_t S + S^2 + J^2 + R(\mathrm{Tr}\, U^2) = 0$,

(2.16) $\qquad \partial_t J + SJ + JS = 0$.

-10-

We deduce from (2.16) that

(2.17)   supp $J(t,\cdot) \subset$ supp J_0 .

Indeed, we can prove (2.17) by noticing that, for any fixed $x \in T^n$

$$\frac{1}{2} \frac{d}{dt} (J(t,x);J(t,x)) = Tr(SJ^2 + JSJ) = 2TrSJ^2 < 2m(t,x)(J(t,x);J(t,x))$$

where $m(t,x)$ is the maximum of the absolute values of the eigenvalues of $S(t,x)$. It follows from Gronwall's inequality that, as long as $S(t,x)$ is smooth, if $J_0(x) = 0$, $J(t,x) = 0$. Let ϕ be an $n \times n$ matrix valued smooth function satisfying the following conditions:

(a)   supp $\phi \cap$ supp $J_0 = \phi$,

(b)   for every $x \in T^n$, $\phi(x)$ is a symmetric, nonnegative matrix, i.e. $\phi(x) = \Psi(x)^2$ for
        some symmetric $\Psi(x)$.

(c) $R\phi = 0$.

(d) $\langle S_0,\phi \rangle < 0$.

Let us postpone for the moment the construction of ϕ and proceed with the proof. Taking the scalar product of (2.15) with ϕ we obtain

(2.18) $\frac{d}{dt} \langle S,\phi \rangle + \langle S^2,\phi \rangle + \langle J^2,\phi \rangle + \langle R(TrU^2),\phi \rangle = 0$.

Now $\langle R(TrU^2),\phi \rangle = \langle TrU^2,R\phi \rangle = 0$ because of assumption (c) and of the symmetry (2.14) of R. Moreover, combining (2.17) and assumption (a) we deduce $\langle J^2,\phi \rangle = 0$. Thus (2.18) becomes

(2.19) $\frac{d}{dt} \langle S,\phi \rangle + \langle S^2,\phi \rangle = 0$.

Now

$$|\langle S,\phi \rangle| = \left| \int_{T^n} Tr(S(x)\phi(x))dx \right| < \int_{T^n} |Tr(S(x)\Psi(x)\Psi(x))|dx <$$

$$< \int_{T^n} (Tr(S(x)\Psi(x))(S(x)\Psi(x))^*)^{1/2}(Tr\Psi(x)\Psi(x))^{1/2}dx <$$

$$< \left(\int_{T^n} TrS(x)\Psi(x)\Psi(x)S(x)dx\right)^{1/2}\left(\int_{T^n} Tr\phi(x)dx\right)^{1/2} = \langle S^2,\phi \rangle^{1/2} \left(\int_{T^n} Tr\phi(x)dx\right)^{1/2} .$$

-11-

It follows from (2.19) that

$$(2.20) \qquad \frac{d}{dt} \langle S, \phi \rangle + \frac{\langle S, \phi \rangle^2}{\int_{T^n} Tr\,\phi(x)\,dx} < 0 .$$

We assumed in (d) that $\langle S_0, \phi \rangle < 0$ and thus we infer that $\langle S(t, \cdot), \phi \rangle$ must become $-\infty$

for t not larger than $T_\infty = \dfrac{\int_{T^n} Tr\,\phi(x)\,dx}{|\langle S_0, \phi \rangle|}$.

We are going to show now how one can construct ϕ satisfying properties (a)-(d).

Let us take a neighborhood V of x_0 such that for $y \in V$, $\Gamma_i(y) \cap supp\ J_0 = \emptyset$. Since

$U_{0ii}(x) = S_{0ii}(x)$ assumption (ii) implies $\int_{\Gamma_i(x_0)} S_{0ii}(x_{01}, \ldots, x, \ldots, x_{0n})\,dx$ is

negative. We may assume that

$$\int_{\Gamma_i(y)} S_{0ii} < 0 \quad \text{for all}\ y \in V .$$

Let $\varphi(x) = \psi(x)^2$ with $\psi(x)$ a smooth function defined in T^n, independent of the i-th

variable (that is, constant on circles $\Gamma_i(z)$ for any $z \in T^n$) with support in

$K = \bigcup_{y \in V} \Gamma_i(y)$ and identically 1 on a set $K_1 = \bigcup_{y \in V_1} \Gamma_i(y)$ for some $x_0 \in V_1 \subset\subset V$. We

define $\phi(x)$ to be the $n \times n$ matrix having all entries equal to zero with the exception

of the entry $\phi(x)_{ii}$ set to be equal to $\varphi(x)$. Clearly properties (a), (b) and (d) are

satisfied from construction. Condition (c) is satisfied for a matrix ϕ if its columns

are divergence free. In the constructed matrix the only nonzero column is the i-th and

$\partial_i \varphi(x) = 0$. This completes the proof of Theorem 2.1.

-12-

3. Solutions with spherical symmetry

In [4] C. Foias showed that the equation (2.10) in the whole space R^n admits solutions of the form $U(t,x) = \beta(t,|x|)(I - n\pi(x))$ where I is the identity matrix and $\pi(x)$ is the projector on the direction x

$$(3.1) \qquad \pi(x) = \left(\frac{x_i x_j}{|x|^2}\right) \qquad i,j = 1,\ldots,n \; .$$

Moreover he obtained a simple equation for the scalar quantity β which blows up. We shall generalize slightly this result, allowing antisymmetric parts in $U(t,x)$. The main reason behind our desire to have nontrivial antisymmetric parts in $U(t,x)$ is that they correspond to the vorticity in the case of genuine Euler equations.

Let A be a rotation, $A \in O(n)$. We denote for a scalar function in R^n, f, by f_A the composed

$$(3.2) \qquad f_A(x) = f(Ax) \; .$$

For a matrix valued function M we denote by M_A the matrix with entries $(M_A)_{ij} = (M_{ij})_A$. We define the operations T_A and $\tilde{T}_A$ on $n \times n$ matrix valued functions as

$$(3.3) \qquad T_A M = A^{-1} M_A A \; ,$$

$$(3.4) \qquad \tilde{T}_A M = (\det A) A^{-1} M_A A \; .$$

Finally, for a matrix valued function U we define $L_A(U)$ by

$$(3.5) \qquad L_A(U) = T_A(S) + \tilde{T}_A(J) \quad \text{where} \quad U = S + J \; ,$$

$S = \frac{1}{2}(U + U^*)$, $J = \frac{1}{2}(U - U^*)$. Let us denote by $N(U)$ the operator giving the distorted Euler equation in R^n:

$$(3.6) \qquad N(U) = \partial_t U + U^2 + R(\mathrm{Tr}\, U^2) \; .$$

Proposition 3.1.

For any $A \in O(n)$, N is covariant with respect to L_A:

$$(3.7) \qquad L_A(N(U)) = N(L_A U) \; .$$

-13-

Corollary 3.2.

If the initial data U_0 is invariant with respect to A, i.e. if $L_A(U_0) = U_0$ then the solution $U(t,x)$ is invariant with respect to A,

$$(3.8) \qquad\qquad L_A(U(t,\cdot)) = U(t,\cdot) \ .$$

Proof of Proposition 3.1.

Let us take the symmetric and antisymmetric parts of $N(U)$

$$(3.9) \qquad\qquad \frac{N(U) + N(U)^*}{2} = \partial_t S + S^2 + J^2 + R(TrU^2) \ ,$$

$$(3.10) \qquad\qquad \frac{N(U) - N(U)^*}{2} = \partial_t J + SJ + JS \ .$$

Applying $\widetilde{T}_A$ to (3.10) we obtain

$$\widetilde{T}_A\left(\frac{N(U) - N(U)^*}{2}\right) = \partial_t \widetilde{T}_A J + (T_A S)(\widetilde{T}_A J) + (\widetilde{T}_A J)(T_A S) = \frac{N(L_A U) - (N(L_A U))^*}{2} \ .$$

In order to check the T_A covariance of the symmetric part of $N(U)$ we make use of the well-known covariance with respect to rotations of the Riesz transforms ([6])

$$(3.11) \qquad\qquad T_A(Rf) = R(f_A)$$

for any scalar function f.

We check now that $Tr(L_A U)^2 = (TrU^2)_A$. Indeed

$$Tr(L_A U)^2 = Tr((T_A S + \widetilde{T}_A J)^2) = Tr((T_A S)^2 + (\widetilde{T}_A J)^2) = Tr(T_A(S^2) + T_A(J^2)) =$$

$$TrA^{-1}(S^2 + J^2)_A A = Tr(S^2 + J^2)_A = (TrU^2)_A \ .$$

Applying T_A to (3.9) we obtain

$$T_A\left(\frac{N(U) + N(U)^*}{2}\right) = \partial_t(T_A S) + (T_A S)^2 + (\widetilde{T}_A J)^2 + R(Tr(L_A U)^2) = \frac{N(L_A U) + (N(L_A U))^*}{2} \ .$$

This proves the proposition. Corollary (3.2) follows from uniqueness of solutions of $N(U) = 0$.

Let us restrict our attention for a moment to the case $n = 3$. Any antisymmetric matrix J defines uniquely a vector $\omega \in \mathbf{R}^3$ such that $Jv = \omega \times v$ for any $v \in \mathbf{R}^3$. Here $\omega \times v$ is the vector $(\omega_2 v_3 - \omega_3 v_2, \omega_3 v_1 - \omega_1 v_3, \omega_1 v_2 - \omega_2 v_1)^t$ and clearly ω is determined by $\omega_1 = J_{32}$, $\omega_2 = J_{13}$, $\omega_3 = J_{21}$. The matrix J^2 can be computed in terms of

-14-

the vector ω:

$$(3.12) \qquad J^2 = -|\omega|^2(I - \pi_\omega)$$

where $\pi_\omega = \dfrac{(\omega_i \omega_j)}{|\omega|^2}$.

We note here that if $J(x)$ is the antisymmetric part of the Jacobian of a function $u(x)$ i.e. $J(x) = \frac{1}{2}(\partial_j u_i - \partial_i u_j)$, $i,j = 1,2,3$ then $\omega(x) = \frac{1}{2}(\nabla \times u)(x)$. If $J(t,x)$ satisfies the antisymmetric part of (2.10) i.e. if

$$(3.13) \qquad \partial_t J + SJ + JS = 0$$

then forming the quantities $\omega(t,x)$ corresponding to $J(t,x)$ we obtain from (3.13) the equation

$$(3.14) \qquad \partial_t \omega = S\omega .$$

This is the analogue of the vorticity equation in the case of Euler equations.

Summarizing, the equation (2.10) is equivalent in the three-dimensional case to

$$(3.15) \qquad \partial_t S + S^2 + J^2 + R(\mathrm{Tr}S^2 + \mathrm{Tr}J^2) = 0$$

coupled with (3.14) where J^2 is given by (3.12). We can consider the system (3.14), (3.15) with J^2 defined by (3.12) in any number of dimensions: S will be a $n \times n$ symmetric matrix and ω an n vector.

Proposition 3.2

Assume that the initial data for the system

$$(3.16) \qquad \partial_t \omega = S\omega ,$$

$$(3.17) \qquad \partial_t S + S^2 + J^2 + R(\mathrm{Tr}(S^2 + J^2)) = 0$$

where $J^2 = -|\omega|^2(I - \pi_\omega)$, $\pi_\omega = \left(\dfrac{\omega_i \omega_j}{|\omega|^2}\right)$, $i,j = 1,\ldots,n$ are of the form

$$(3.18) \qquad \omega_0(x) = \gamma_0(|x|)\cdot \frac{x}{|x|} , \qquad x \in \mathbb{R}^n ,$$

$$(3.19) \qquad S_0(x) = \beta_0(|x|)(I - n\pi(x)), \ \pi(x) \text{ given in (3.1).}$$

Then for as long as the solution $S(t,x)$, $\omega(t,x)$ stays smooth, it has the form

$$(3.20) \qquad \omega(t,x) = \gamma(t,|x|) \frac{x}{|x|} ,$$

-15-

(3.21) $$S(t,x) = \beta(t,|x|)(I - n\pi(x))$$

where γ, β are two scalar functions satisfying

(3.22) $$\partial_t\beta + \beta^2 - \gamma^2 - \frac{n-1}{r^n}\int_0^r s^{n-1}(n\beta^2 - \gamma^2)ds = 0 \ ,$$

(3.23) $$\partial_t\gamma + (n-1)\gamma\beta = 0 \ ,$$

(3.24) $$\gamma(0,r) = \gamma_0(r), \quad \beta(0,r) = \beta_0(r) \ ,$$

(3.25) $$\gamma(t,0) = \gamma(t,\infty) = 0, \quad \beta(t,0) = \beta(t,\infty) = 0 \ .$$

Remark 1

The equation obtained by Foias is the particular case $\gamma(t,r) \equiv 0$ arising from $\gamma_0 \equiv 0$.

Remark 2

In $n = 3$ initial data of the form (3.18), (3.19) are those which satisfy $L_A U_0 = U_0$, for all $A \in O(3)$, $Tr U_0 = 0$.

We start by computing Rf for a radial function.

Lemma 3.3

Let $f = f(r)$ be a smooth function defined for $r > 0$ decaying sufficiently at infinity (for instance $f(|x|) \in L^1 \cap L^2$ in R^n). Then

(3.26) $$Rf = -[gI + k\pi]$$

when g and k are radial functions defined by

(3.27) $$g(r) = \frac{1}{r^n}\int_0^r s^{n-1}f(s)ds$$

(3.28) $$k(r) = f(r) - \frac{n}{r^n}\int_0^r s^{n-1}f(s)ds \ .$$

Proof

Let us use the notation $f' = \dfrac{df}{dr}$, $r = |x|$. Then
$\partial_{ij}^2 f(r) = \dfrac{f'(r)}{r}\delta_{ij} + \left(\left(\dfrac{f'(r)}{r}\right)' r\right)\pi(x)$. On the other hand

-16-

$$\Delta(g(r)\delta_{ij} + h(r)x_ix_j) = (\Delta g + 2h)\delta_{ij} + (\Delta h + \frac{4h'}{r})x_ix_j \ .$$

Thus $\partial^2_{ij}f(r) = \Delta(g(r)\delta_{ij} + h(r)x_ix_j)$ if the system

(3.29)
$$\Delta g + 2h = \frac{f'(r)}{r} \ ,$$

(3.30)
$$r\Big(\Delta h + \frac{4h'}{r}\Big) = \Big(\frac{f'(r)}{r}\Big)' \quad \text{is solved} \ .$$

Now (3.30) follows from (3.29) if

$$(\Delta g)' + 2h' = r\Big(\Delta h + \frac{4h'}{r}\Big) \quad \text{i.e. if} \quad (\Delta g)' = (rh')' + nh' \ .$$

This follows if $\Delta g = (rh)' + \frac{n-1}{r}(rh)$. So (3.30) is a consequence of (3.29) if

$g' = rh$. With this choice for g we solve (3.29):

$$(rh)' + \frac{n-1}{r}\,rh + 2h = \frac{f'}{r} \ .$$

This gives $(r^nk)' = r^nf'$ for $k = r^2h$. We obtain the formula (3.28) for k:

(3.31)
$$k(r) = f(r) - \frac{n}{r^n}\int_0^r s^{n-1}f(s)ds \ .$$

Then $g' = \frac{1}{r}k$. In order to check (3.27) let us note that

$$(ng + k - f)' = \frac{n}{r}k + k' - f' = 0 \ .$$

Thus, since all these functions vanish at infinity we obtain

(3.32)
$$f = ng + k \ .$$

Therefore (3.31) and (3.32) imply (3.27). We note that (3.32) follows also from the

familiar $R_iR_if = -f$ (see [6]) by taking the trace in (3.26).

<u>Proof of Proposition 3.2</u>

We shall use the ansatz $\omega(t,x) = \gamma(t,|x|)\frac{x}{|x|}$, $S(t,x) = \beta(t,|x|)(I - n\pi(x))$ and

check that equations (3.16), (3.17) give consistent equations for β, γ. Equation (3.16)

becomes $\partial_t\gamma = (1-n)\beta\gamma$ i.e. (3.23). Now $S^2 = \beta^2(I - n\pi)^2 = \beta^2(I + (n^2 - 2n)\pi)$

because $\pi^2 = \pi$. Also $J^2 = -\gamma^2(I - \pi)$. Indeed $\pi_\omega = \pi(x)$ because x and ω define

the same direction. In order to proceed we put $f(r) = \text{Tr}(S^2 + J^2)$ and compute

(3.33)
$$\text{Tr}(S^2 + J^2) = f(r) = (n-1)[n\beta^2 - \gamma^2] \ .$$

According to Lemma 3.3 it follows that

-17-

(3.34)
$$R(\operatorname{Tr}(S^2 + J^2)) = -(gI + k\pi)$$

with g,k defined by (3.27), (3.28) and f by (3.33). At this point equation (3.17) has the form

(3.35)
$$(\partial_t \beta)(I - n\pi) + (\beta^2 - \gamma^2 - g)I + ((n^2 - 2n)\beta^2 + \gamma^2 - k)\pi = 0 .$$

The only way in which (3.35) can possibly give a consistent equation for β is if it factors out $(I - n\pi)$, that is if

(3.36)
$$(n^2 - 2n)\beta^2 + \gamma^2 - k = -n(\beta^2 - \gamma^2 - g) .$$

But (3.36) is equivalent to

$$(n - 1)(n\beta^2 - \gamma^2) = ng + k$$

which in view of (3.33) is nothing but (3.32) in disguise. Therefore the equation (3.35) becomes

(3.37)
$$(\partial_t \beta + \beta^2 - \gamma^2 - g)(I - n\pi) = 0$$

which is satisfied if β solves (3.22) because (3.27) and (3.33) imply

$$g(r) = \frac{n - 1}{r^n} \int_0^r s^{n-1}(n\beta^2 - \gamma^2)ds.$$

We present now the blow-up argument.

Theorem 3.4

Let us assume that beside the conditions (3.18) and (3.19) of Proposition 3.2 being fulfilled, the initial data for the system (3.16), (3.17) satisfy also

(3.38)
$$\gamma_0(r) = 0 \quad \text{for} \quad 0 < r < R_1 \quad \text{for some} \quad R_1 > 0 ,$$

(3.39)
$$\beta_0(r) = 0 \quad \text{for} \quad 0 < r < R \quad \text{for some} \quad 0 < R < R_1 ,$$

(3.40)
$$\int_R^{\alpha R} \beta_0(r)r^{n-1}dr < 0 \quad \text{for some} \quad 1 < \alpha, \ (\alpha \ \text{near} \ 1) .$$

Then the solution to the distorted Euler equations (3.16), (3.17) having $\omega_0(x) = \gamma_0(|x|) \dfrac{x}{|x|}$, $S_0(x) = \beta_0(|x|)(I - n\pi(x))$ for initial data breakdown in finite time. More precisely S breaks down near $|x| = R$.

-18-

Proof

As proven in Proposition (3.2) the solutions are $\omega(t,x) = \gamma(t,|x|)\frac{x}{|x|}$,
$S(t,x) = \beta(t,|x|)(I - n\pi)$ with γ,ω satisfying (3.22)-(3.25). Since (3.23) can be integrated

$$(3.41) \qquad\qquad \gamma(t,r) = \gamma_0(r)\exp \int_0^t (1 - n)\beta(s,r)ds$$

it follows from (3.38) that

$$(3.42) \qquad\qquad \gamma(t,r) = 0 \quad\text{for}\quad 0 < r < R_1 .$$

Therefore, for $0 < r < R_1$ equation (3.22) becomes

$$(3.43) \qquad\qquad \partial_t\beta + \beta^2 - \frac{n - 1}{r^n} \int_0^r s^{n-1}n\beta^2(s)ds = 0 .$$

Now we claim that from (3.43) it follows that property (3.39) is preserved by $\beta(t,r)$, $t > 0$, as long as both β and γ are smooth. Indeed multiplying (3.43) by $\beta(t,r)r^{n-1}$ and integrating between 0 and R one obtains

$$\frac{1}{2}\frac{d}{dt}\int_0^R \beta^2(t,r)r^{n-1}dr + \int_0^R \beta^3(t,r)r^{n-1}dr = (n - 1)n \int_0^R \beta^2(t,\rho)\rho^{n-1} \int_\rho^R \frac{\beta(t,r)}{r} dr \quad\text{or}$$

$$\frac{1}{2}\frac{d}{dt}\int_0^R \beta^2(t,r)r^{n-1}dr = \int_0^R \beta^2(t,r)r^{n-1}[n(n - 1) \int_r^R \frac{\beta(t,\rho)}{\rho} d\rho - \beta(t,r]dr .$$

Thus

$$\frac{d}{dt}\int_0^R \beta^2(t,r)r^{n-1}dr < 2 \max_{\rho\in[0,R]} (n(n - 1) \int_\rho^R \frac{\beta(t,s)}{s} ds - \beta(t,\rho))\int_0^R \beta^2(t,r)r^{n-1}dr$$

and therefore, as long as β is smooth

$$\int_0^R \beta^2(t,r)r^{n-1}dr < \int_0^R \beta_0^2(r)r^{n-1}dr \exp 2 \int_0^t \max_{\rho\in[0,R]} (n(n - 1) \int_\rho^R \frac{\beta(\tau,s)}{s} ds - \beta(\tau,\rho))d\tau$$

-19-

and by (3.39) it follows

$$(3.44) \qquad \int_0^R \beta^2(t,r)r^{n-1}dr = 0, \qquad t > 0 .$$

Let us take now $\alpha > 1$ such that (3.40) is valid and α small enough such that $\alpha R < R_1$, $n(n-1)\log \alpha < 1$. Integrating in (3.43) between R and αR we obtain

$$(3.45) \qquad \frac{d}{dt} \int_R^{\alpha R} \beta(t,r)r^{n-1}dr + (1 - n(n-1)\log \alpha) \int_R^{\alpha R} \beta^2(t,s)s^{n-1}ds <$$

$$< n(n-1)\log \alpha \int_0^R \beta^2(t,s)s^{n-1}ds = 0 .$$

Now

$$\int_R^{\alpha R} \beta^2(t,s)s^{n-1}ds > \frac{n}{R^n} \frac{1}{\alpha^n - 1} (\int_R^{\alpha R} \beta(t,s)s^{n-1}ds)^2 .$$

It follows that

$$(3.46) \qquad \frac{d}{dt} \int_R^{\alpha R} \beta(t,r)r^{n-1}dr + \frac{n(1 - n(n-1)\log \alpha)}{R^n(\alpha^n - 1)} (\int_R^{\alpha R} \beta(t,s)s^{n-1}ds)^2 < 0$$

and since

$$\int_R^{\alpha R} \beta_0(r)r^{n-1}dr < 0$$

we conclude that $\int_R^{\alpha R} \beta(t,r)r^{n-1}dr$ becomes $-\infty$ for t not larger than

$$T_\infty = \frac{R^n(\alpha^n - 1)}{n(1 - n(n-1)\log \alpha)} \cdot \frac{1}{|\int_R^{\alpha R} \beta_0(r)r^{n-1}dr|} .$$

-20-

One can easily obtain a blow-up argument for equations (3.22), (3.23) at the origin if one drops the requirement that $\beta_0(0) = 0$. However, this would lead to functions $S(t,x)$ which are not defined at $x = 0$.

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We prove that as long as the solution to the ideal flow equation in three space dimensions is smooth the solutions of slightly viscous incompressible Navier-Stokes equations are smooth. We describe blow-up results for a semi-Lagrangian short time approximation of the incompressible Euler equation.

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